n = 5

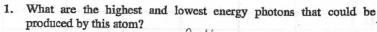
n = 4

-1.9 eV

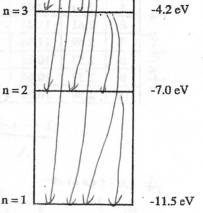
-2.8 éV

QUESTIONS AND PROBLEMS SPECTROSCOPY

Use the energy level diagram to the right to answer the following four questions:



10



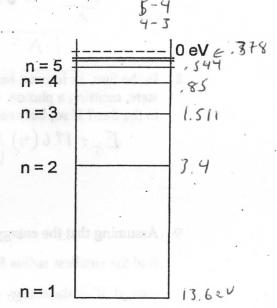
3. The range of wavelengths for light is roughly:

$<4\times10^{-7}m \Rightarrow ultraviolet$	Energy rapiges
$4 \times 10^{-7} m - 5 \times 10^{-7} m \Rightarrow blue$	2.48 < E < 3.1
$5 \times 10^{-7} m - 6 \times 10^{-7} m \Rightarrow green$	2.067 < Ex < 2.48
$6 \times 10^{-7} m - 7 \times 10^{-7} m \Rightarrow red$	1.77 < Ex 22.067
$> 7 \times 10^{-7} m \Rightarrow infrared$	Ex < 1.77

Using this information, determine which of the possible transitions leads to photons that are infrared, visible, or ultraviolet. If they are visible, list their color.

To the right is the energy-level diagram for Hydrogen. Lots of scientists have studied the hydrogen spectrum because hydrogen is the most abundant element in the universe. All stars mostly consist of Hydrogen atoms. The hydrogen spectrum is divided into different series, named after the men who studied them.

- Lyman Series: All frequencies where the electron drops to n = 1 state.
- Balmer Series: All frequencies where the electron drops to n = 2 state.
- Paschen Series: All frequencies where the electron drops to n = 3 state.



5. Find the color of the 4 visible spectral lines that constitute the Balmer Series.

6. Below is a set of data from an unknown atom. The wavelengths of spectral lines are given in nm. The rows are based on starting energy states and the columns are the ending energy states. Use these wavelengths to identify the element.

	Ending E-States										
		1	2	. 3	4	5	6	7	8	9	
States	1		- 1 t-	n mol	garmollait	nd and	in the ed	OF HERMAN	a bornl van	pro buil set	
	2	3.38	4.1							ALC: USON	
	3	2.85	18.25	0	describer			1 7		s unit.	
	4	2.70	13.52	52.14				late	a sistema t	n in the second	
1	5	2.64	12.07	35.64	112.64		Variation of	2 5			
	6	2.61	11.40	30.41	72.99	207.36	10				
(+17)	7	2.59	11.04	27.94	60.21	129.36	343.89	P. Day	Abdis		
2	. 8	2.57	10.81	26.54	54.07	103.97	208.54	529.85			
S E	9	2.57	10.66	25.66	50.53	91.64	(164.23)	314.34	772.83	family .	
2	10	2.56	10.56	(25.07)	48.27	84.48	142.56	243.50	450.55	1080.44	

A:
$$\lambda = 60$$
, $z = 13$, 60 z^{2} $(\frac{1}{h_{1}}z - \frac{1}{h_{2}})$

A: $\lambda = 60$, $z = 10$
 $n_{1} = 7$
 $n_{2} = 7$
 $n_{3} = 7$
 $n_{4} = 7$
 $n_{5} = 6$
 $n_{7} = 7$
 $n_{7} = 9$
 $n_{$

7. Elements in space can absorb radiation to excite their electrons to higher energy states. This is called the absorption spectrum. If an electron is in the 2nd energy state of Beryllium (4 protons) what are the 3 longest wavelength of light that the gas can absorb?

$$\frac{1240}{\lambda} = 13.6(4^2)\left(\frac{1}{2^2} - \frac{1}{n^2}\right)$$

$$1 = 41.03nn, 30.4nn, 27.14nn$$

8. In the Sun, an ionized helium (He⁺) atom makes a transition from the 6th to the 2nd state, emitting a photon. Can the photon be absorbed by hydrogen atoms present in the Sun? If so, between what energy states will the hydrogen atom jump?

9. Assuming that the energy of the electron is all electric potential energy $U_{elec} = \frac{kqq_2}{r}$ find the smallest radius for an electron in a hydrogen atom (known as the Bohr radius). If all the energy were gravitational energy $U_{grav} = \frac{6.67E - 11m_1m_2}{r}$ what would be the Bohr radius now? Hint: me = 9.11*10⁻³¹ kg mp = 1.6726*10⁻²⁷kg

$$E := [3.6eV = -2.18.15^{18}J = \frac{ke^2}{r} \int_{-2.18\cdot15^{18}J} = \frac{(6.17\cdot15^{18})(9.11\cdot15^{31})(1.67\cdot15^{17})}{r}$$

$$-2.18\cdot15^{18}J = \frac{(6.17\cdot15^{11})(9.11\cdot15^{31})(1.67\cdot15^{17})}{r}$$